

March 11, 2016 NEXT meeting (50 min including discussion)  
Kyoto

# Threat of the Sun and Superflare

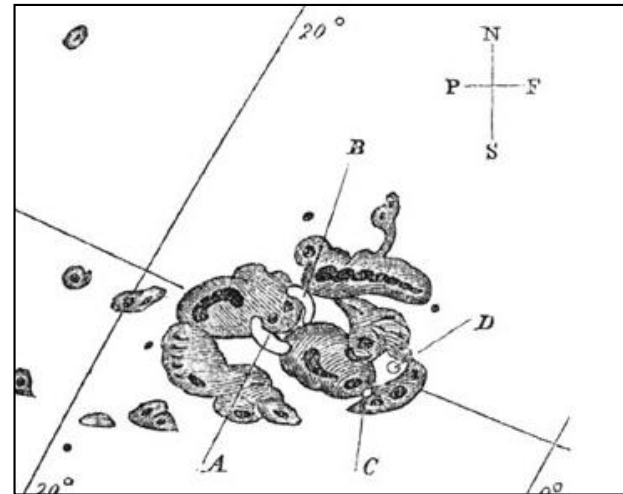
Kazunari Shibata

Kwasan and Hida observatories,  
Kyoto University

# Carrington flare

(1859, Sep 1, am 11:18 )

- The **first flare** that human beings observed
- by Richard Carrington (England)
- white flare for 5 minutes
- **very bright aurora** appeared next day morning at many places on Earth, e.g. Cuba, the Bahamas, Jamaica, El Salvador, and Hawaii.
- Largest magnetic storm (> 1000 nT) in recent 200 yrs.



**Telegraph systems all over Europe and North America failed.**

**Telegraph pylons threw sparks and telegraph paper spontaneously caught Fire (Loomis 1861)**

# Magnetic storm and aurora on 1989 March 13, that lead to Quebec blackout

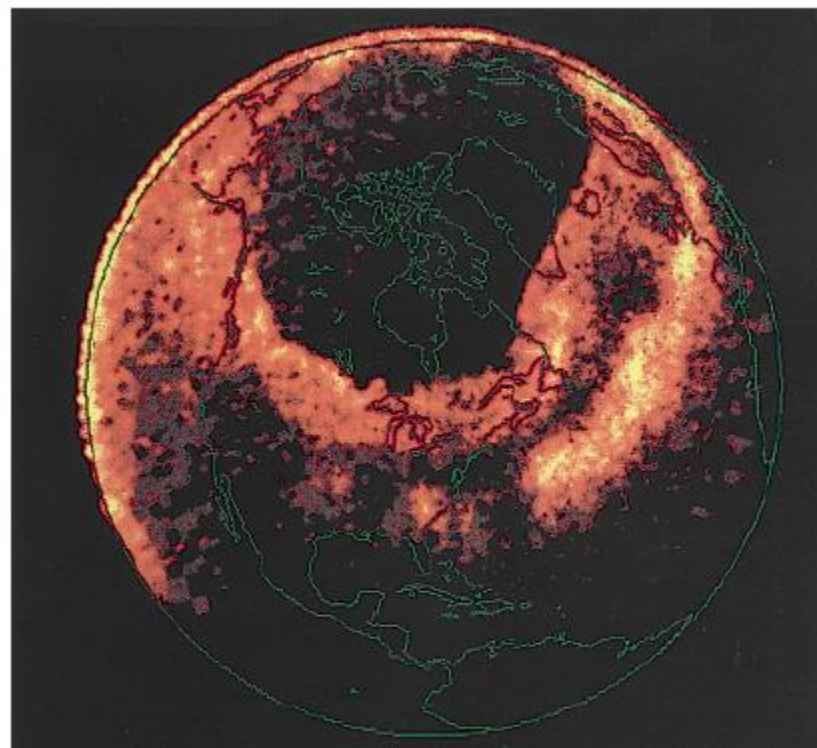
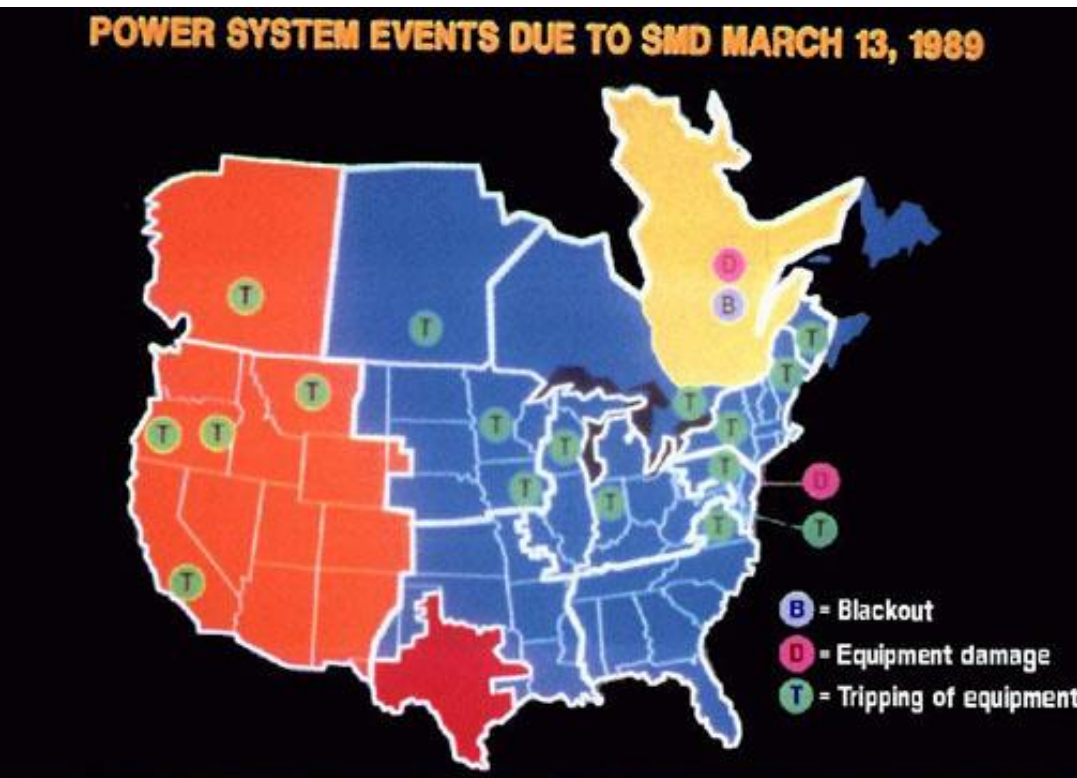


図4 発電所の事故があった日に人工衛星で撮影されたオーロラ。カナダ一帯に強いオーロラが現れたのがわかる。(アイオワ大学 L. A. Frank 教授)

Magnetic storm  $\sim 540$  nT

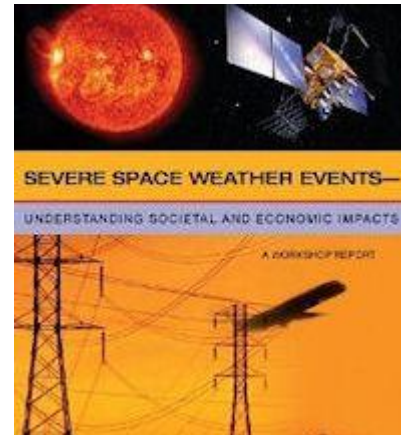
Produced by a big solar flare (X-class: X4.6)

<http://www.stelab.nagoya-u.ac.jp/ste-www1/pub/ste-nl/Newsletter28.pdf>



# Will the Carrington-class flare occur again ?

- If the Carrington-class flare occur now, what will happen ?
- According to a study by the National Academy of Sciences (2008), the total economic impact could exceed **\$2 trillion**



[http://www.nap.edu/catalog.php?record\\_id=12507](http://www.nap.edu/catalog.php?record_id=12507)

Will the Carrington-class flare  
occur again ?

Can much bigger flares, **superflares**  
( $>10^{33}$  erg), occur on the Sun at present ?

What is the impact of superflares on the  
Earth, if **superflares** would occur on the  
Sun ?

To answer these questions  
is the subject of my talk.

# contents

- Introduction
- Solar Flares
- Superflares on Solar Type Stars
- Can Superflares Occur on Our Sun ?

# Solar Flares

# Solar flare

Discovered in 19c

Explosive energy release

That occur near sunspot

**magnetic energy is the  
source of energy**

Size  $\sim 10^9 - 10^{10}$  cm

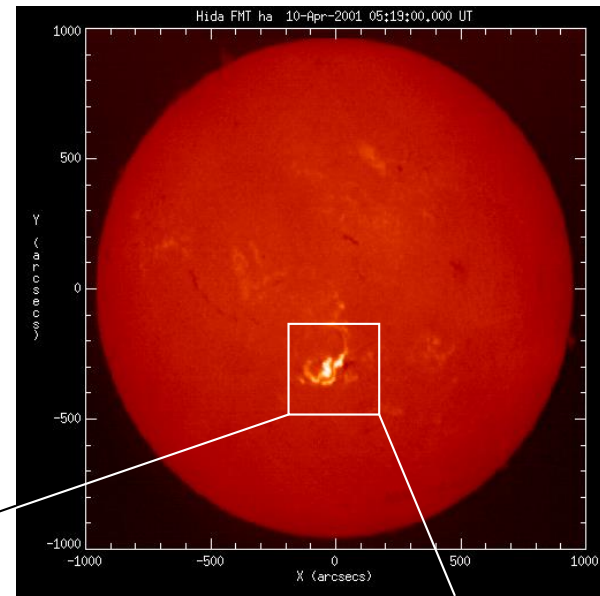
Time scale  $\sim 1\text{min} - 1\text{hour}$

Total energy  $\sim$

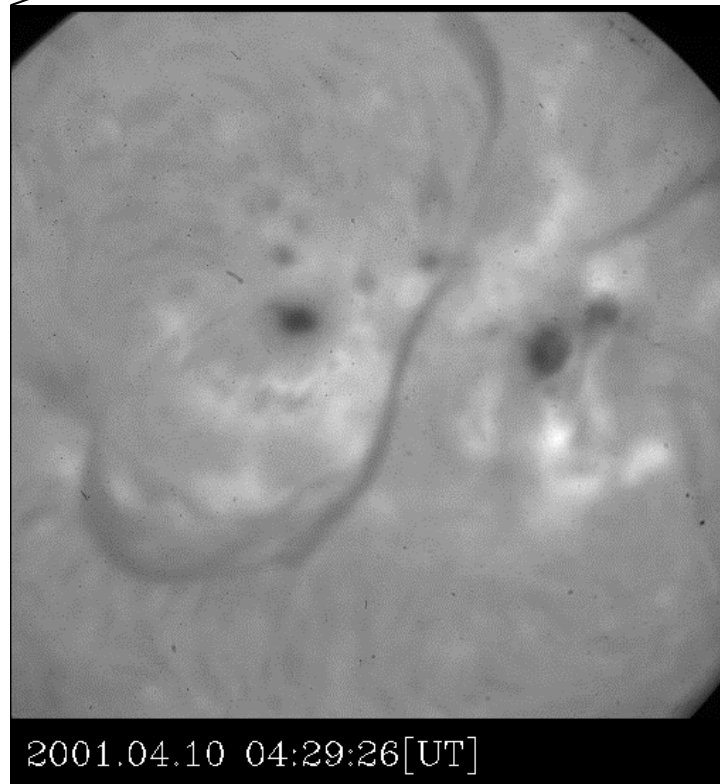
$10^{29} - 10^{32}$  erg

Mechanism has been  
puzzling since 19c until  
recently

Hida Obs/Kyoto U.



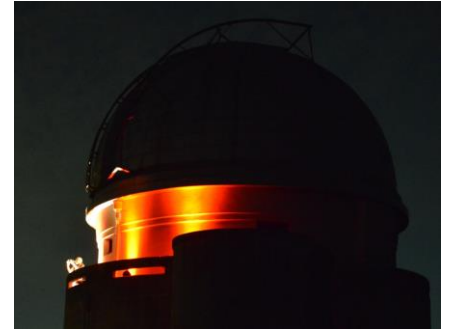
H $\alpha$





# Special entertainment

## Kojiki and Universe



Shibata and Kitaro (musician, Grammy Award Winner)

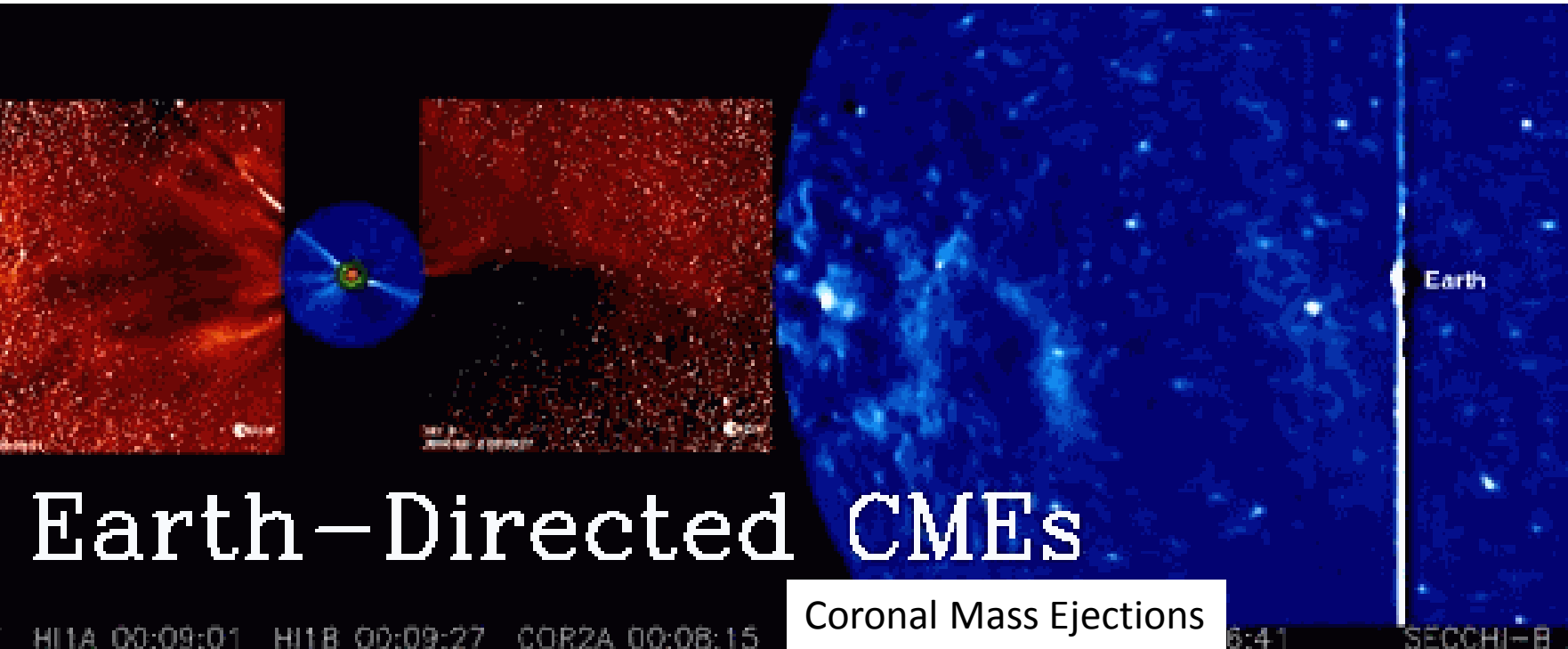
Let's enjoy Kitaro-san's  
Music Kojiki (chap 4:  
Orochi = Monster) with  
beautiful movies of  
**solar flares and  
eruptions**

(7 min)



# Blast Waves from the Sun to the Earth Observed with STEREO

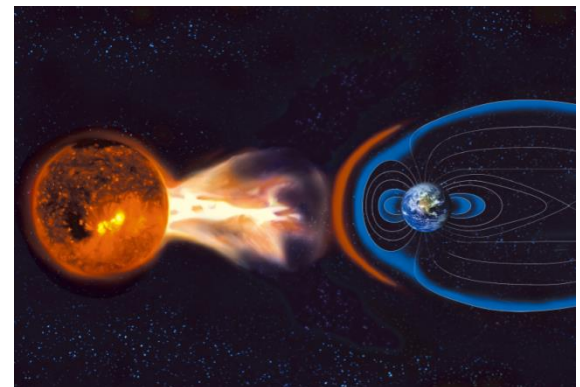
Courtesy of M. Temmer



← 1AU=1.5x10<sup>8</sup> km →

These phenomena lead to magnetic storm and various hazards on the Earth

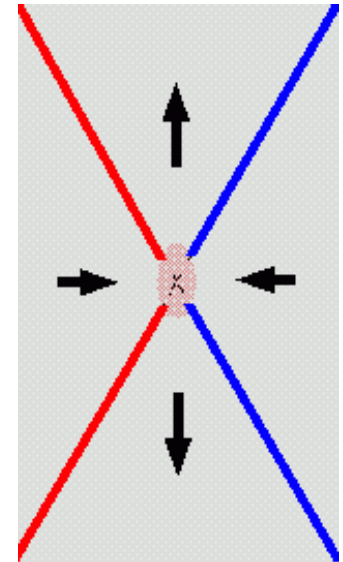
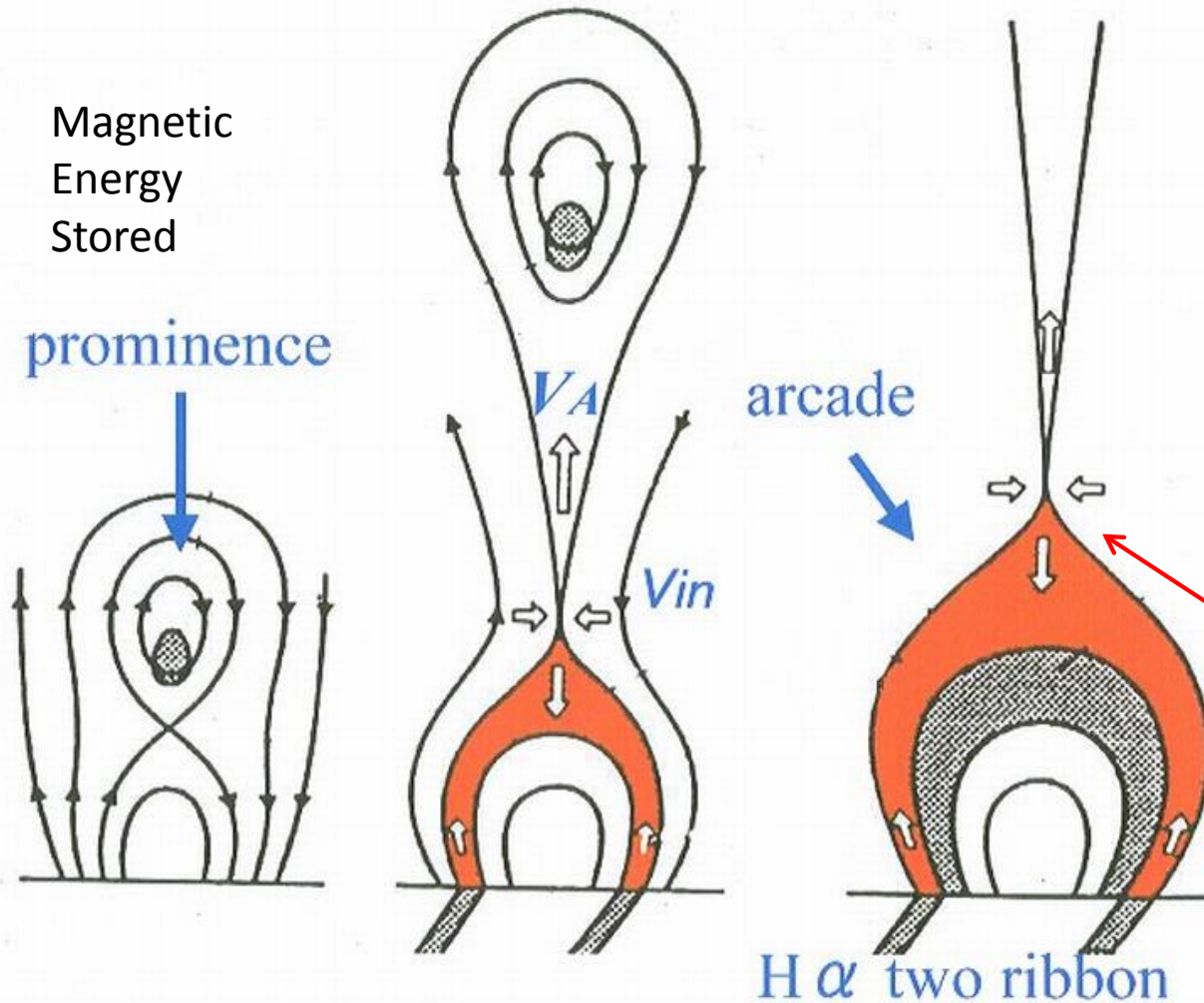
=> **Space weather prediction** is an urgent issue



What is the Mechanism of  
Solar Flares ?

# Standard Flare Model

(Carmichael 1964, Sturrock 1966, Hirayama 1974,  
Kopp and Pneuman 1976)



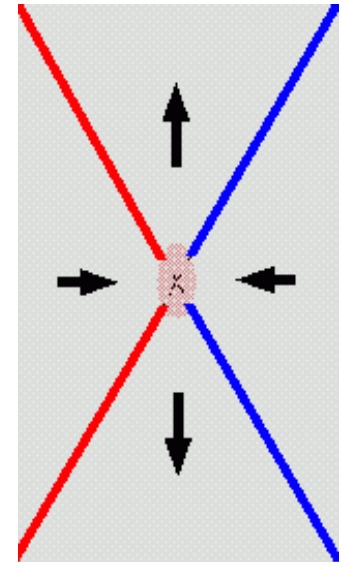
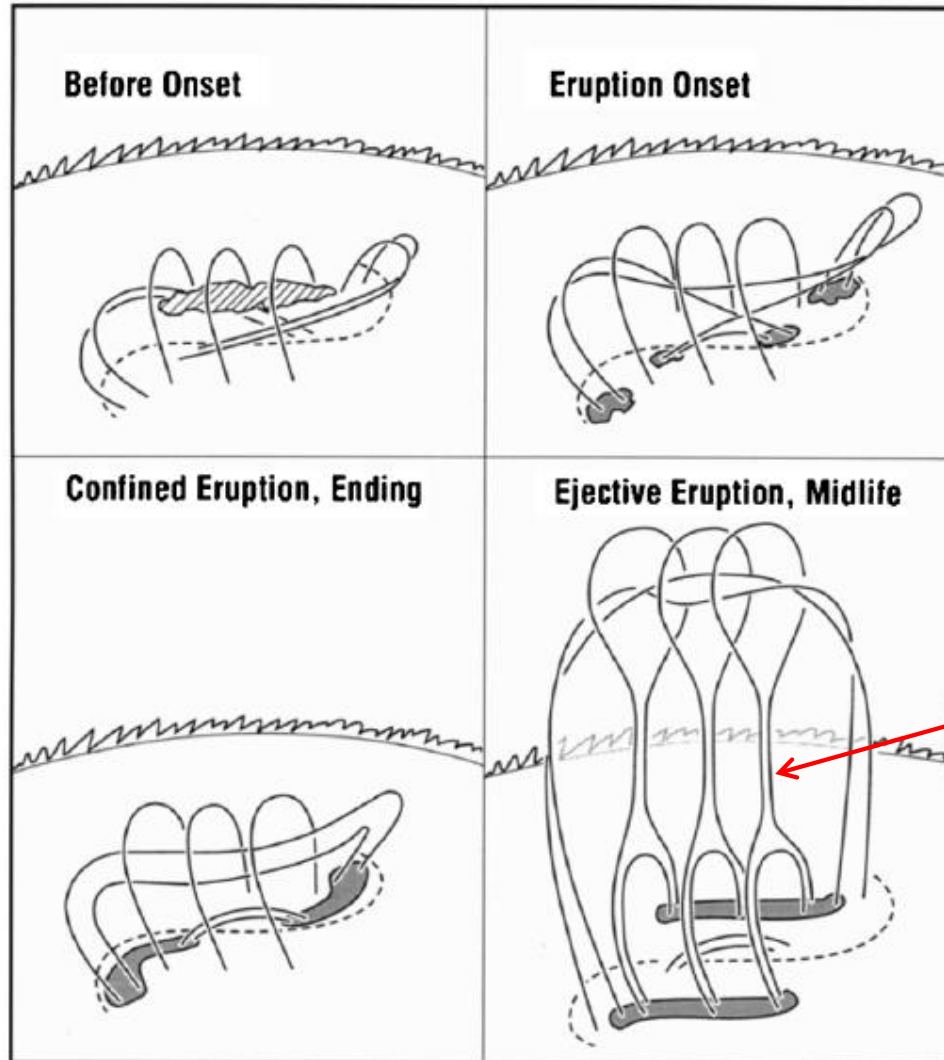
Courtesy of E. Priest

**Magnetic  
Reconnection**

# Standard Flare Model

(Carmichael 1964, Sturrock 1966, Hirayama 1974,  
Kopp and Pneuman 1976)

Magnetic  
Energy  
Stored



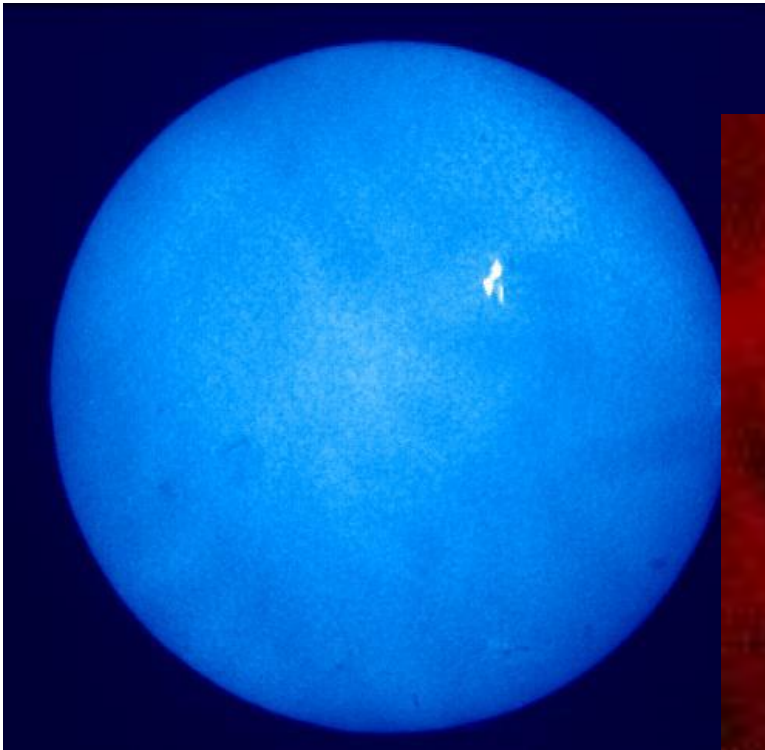
Courtesy of E. Priest

**Magnetic  
Reconnection**

From Moore 2001

# Simultaneous H $\alpha$ and X-ray view of a flare

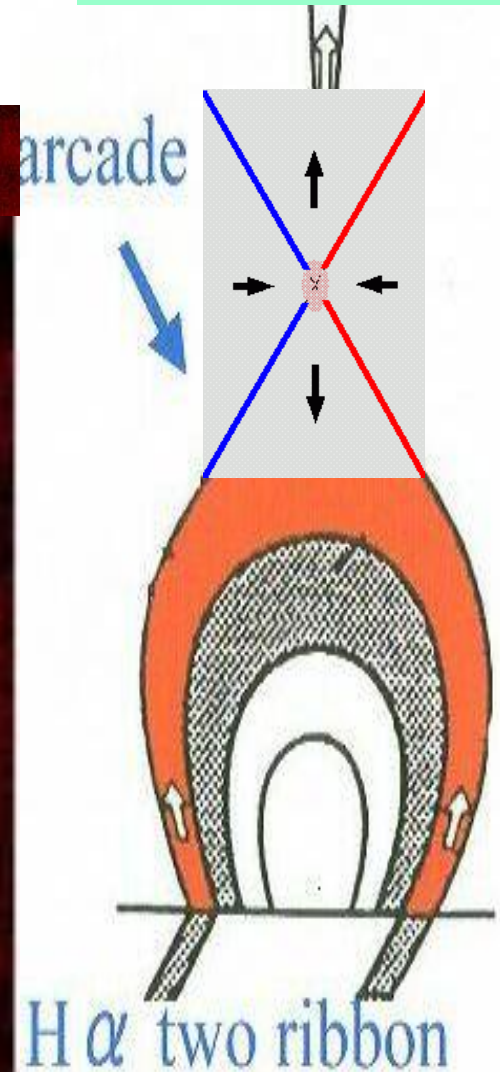
H $\alpha$



X-ray



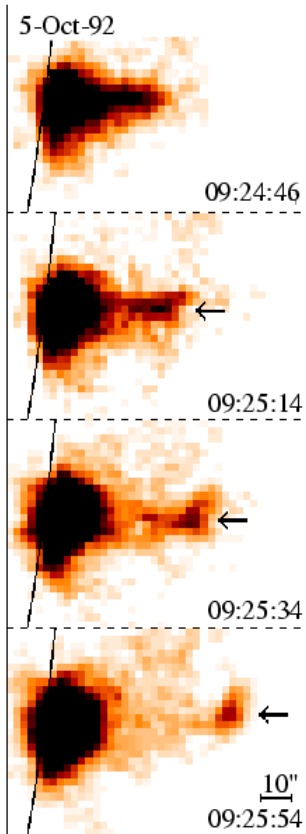
**Magnetic reconnection**



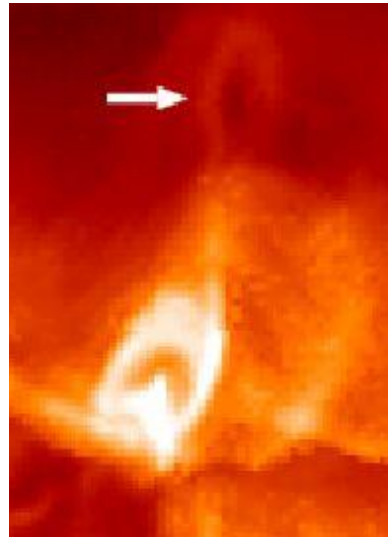
arcade

H $\alpha$  two ribbon

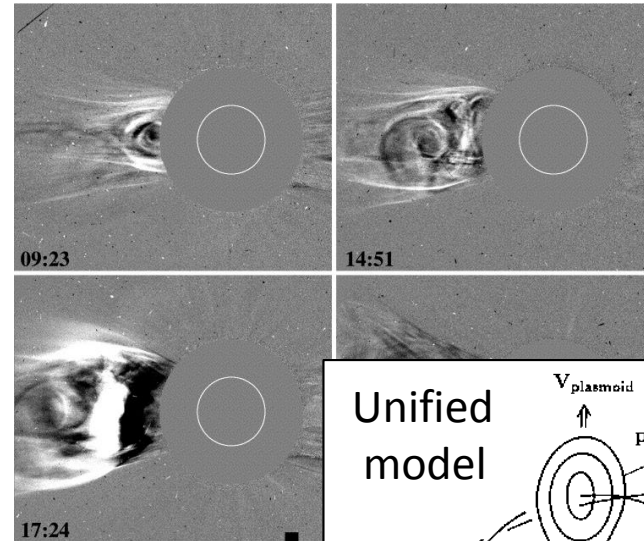
# Plasmoid ejections are ubiquitous



impulsive flares  
 $\sim 10^9$  cm  
 (Ohyama+S 1998)

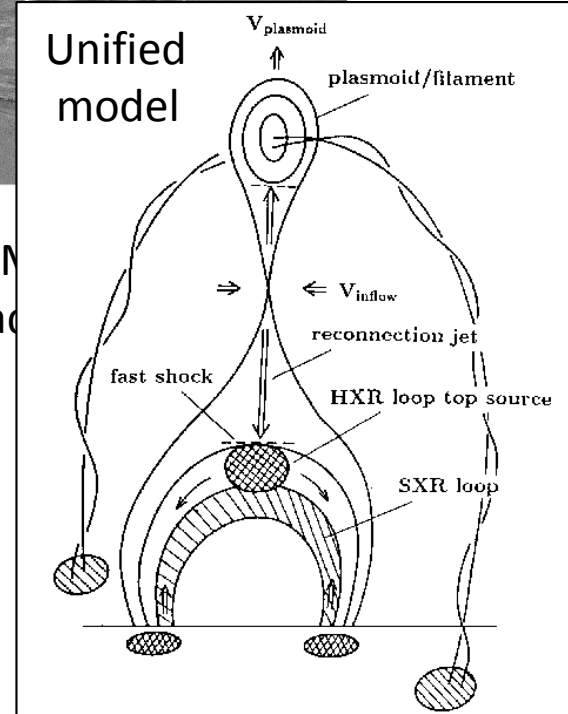


LDE(Long Duration Event) flares  
 $\sim 10^{10}$  cm  
 (Tsuneta 1992, Hudson 1993)



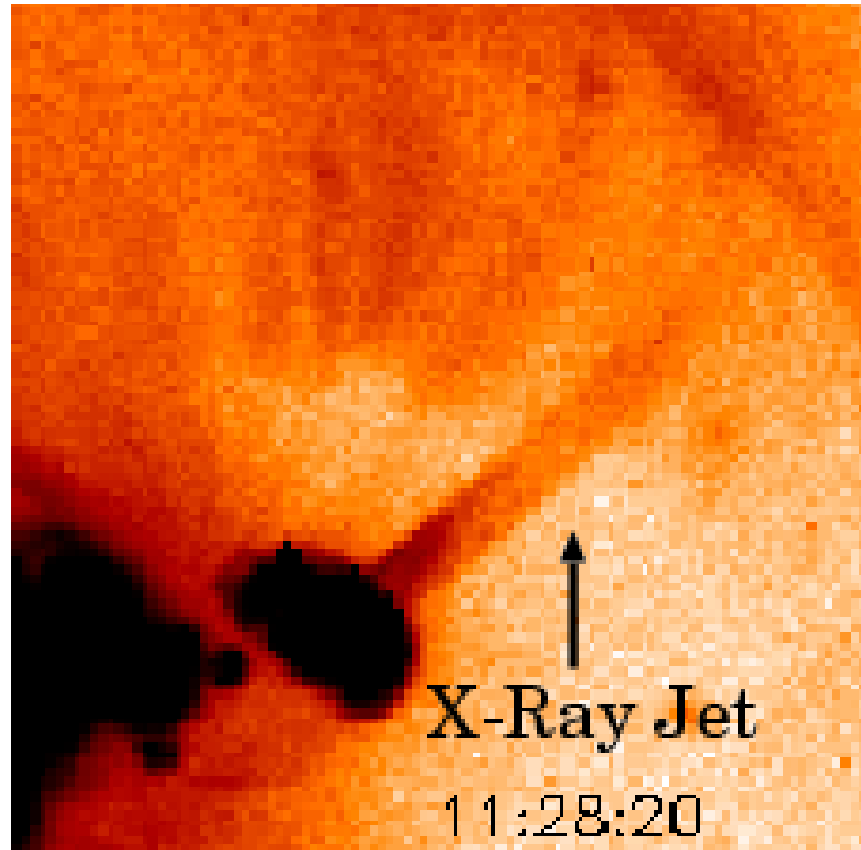
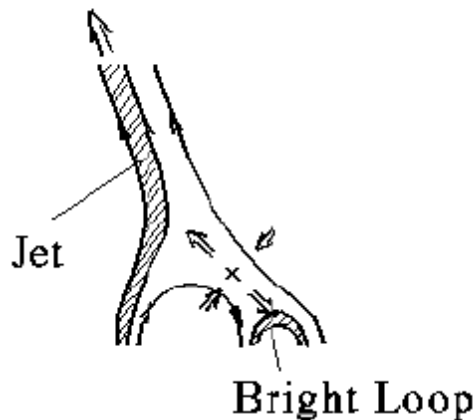
CMEs(Coronal Mass Ejections)  
 from Giant arcades  
 $\sim 10^{11}$  cm  
 (Dere 1995)

Plasmoid-Induced-Reconnection  
 (Shibata 1999)



# Jets from very small flares (microflares)

- Yohkoh/SXT discovered X-ray jets from microflares (Shibata et al. 1992, Strong et al. 1992, Shimojo et al. 1996)





$$t_A = L / V_A$$

$$V_A = \frac{B}{\sqrt{4\pi\rho}}$$

(Alfven speed)

# Summary of observations of “flares” in the solar atmosphere

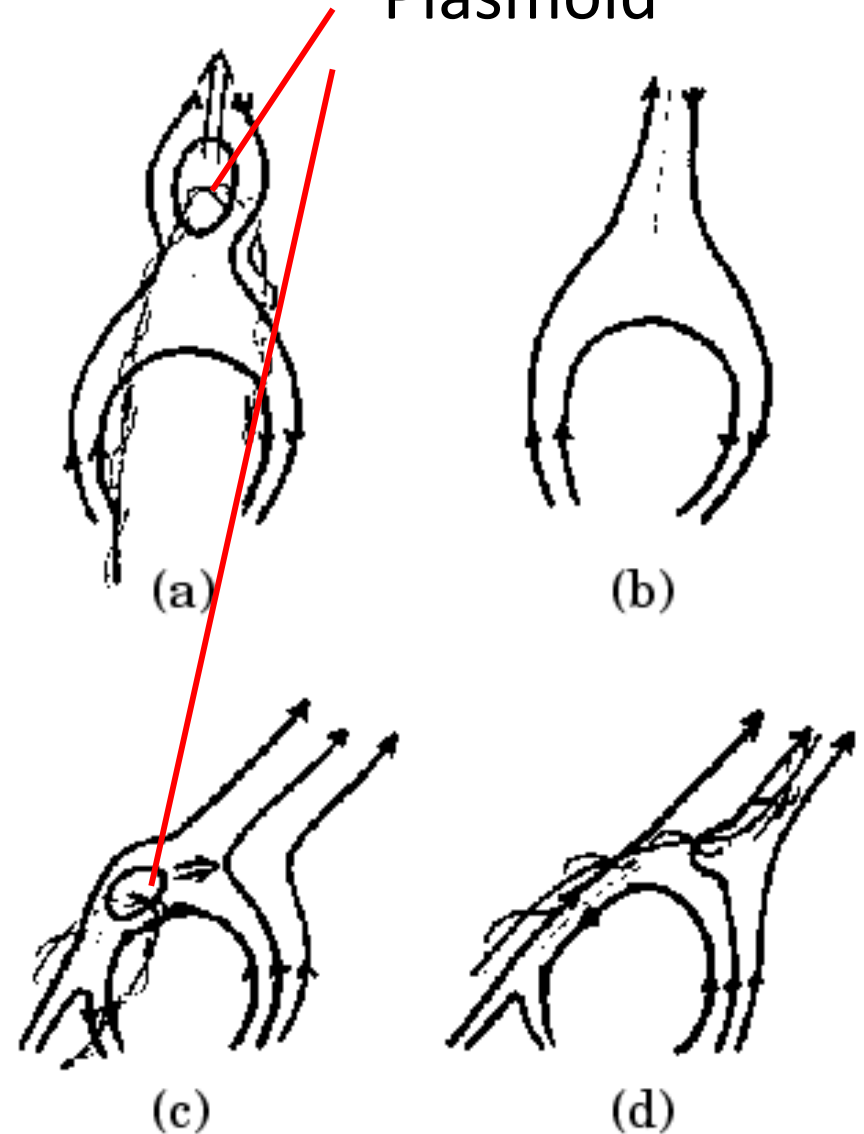
“flares”	Size (L)	Lifetime (t)	Alfven time ( $t_A$ )	$t/t_A$	Mass ejection
Microflares	$10^3 - 10^4$ km	100-1000sec	1-10 sec	~100	jet/surge
Impulsive flares	$(1-3) \times 10^4$ km	10 min – 1 hr	10-30 sec	~60-100	X-ray plasmoid/ Spray
Long duration (LDE) flares	$(3-10) \times 10^4$ km	1-10 hr	30-100 sec	~100-300	X-ray plasmoid/ prom. eruption
Giant arcades	$10^5 - 10^6$ km	10 hr – 2 days	100-1000 sec	~100-300	CME/prom. eruption

**Unified model**  
 (plasmoid-induced  
 reconnection model)  
 (Shibata 1996, 1999,  
 Shibata and Tanuma 2001)

(a,b): large scale flares,  
 Coronal mass ejections

(c,d) : small scale flares,  
 microflares, jets

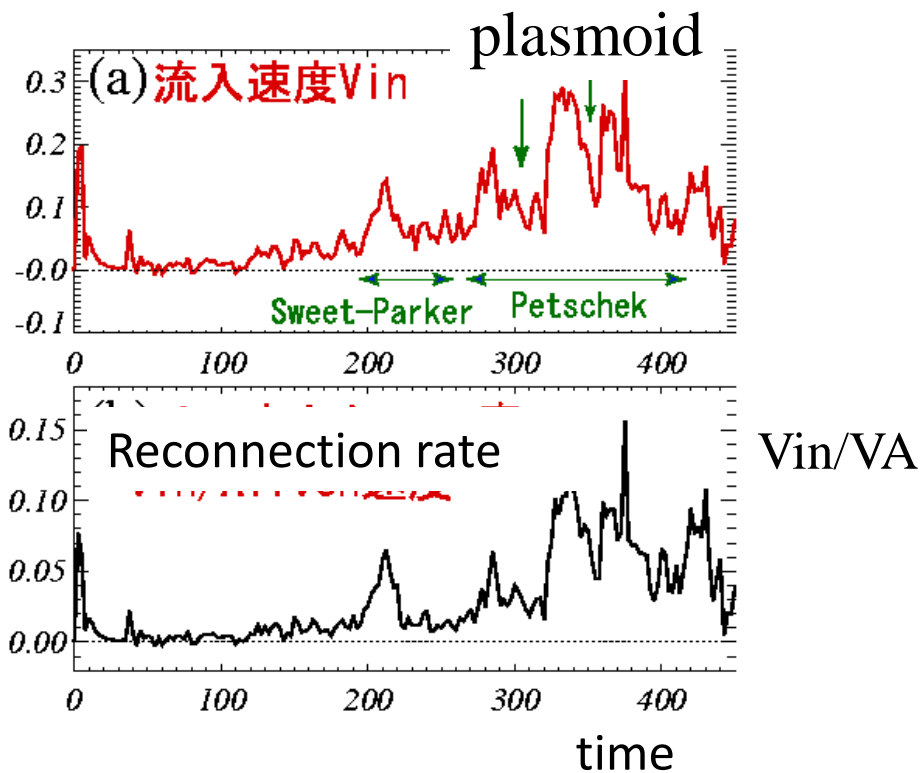
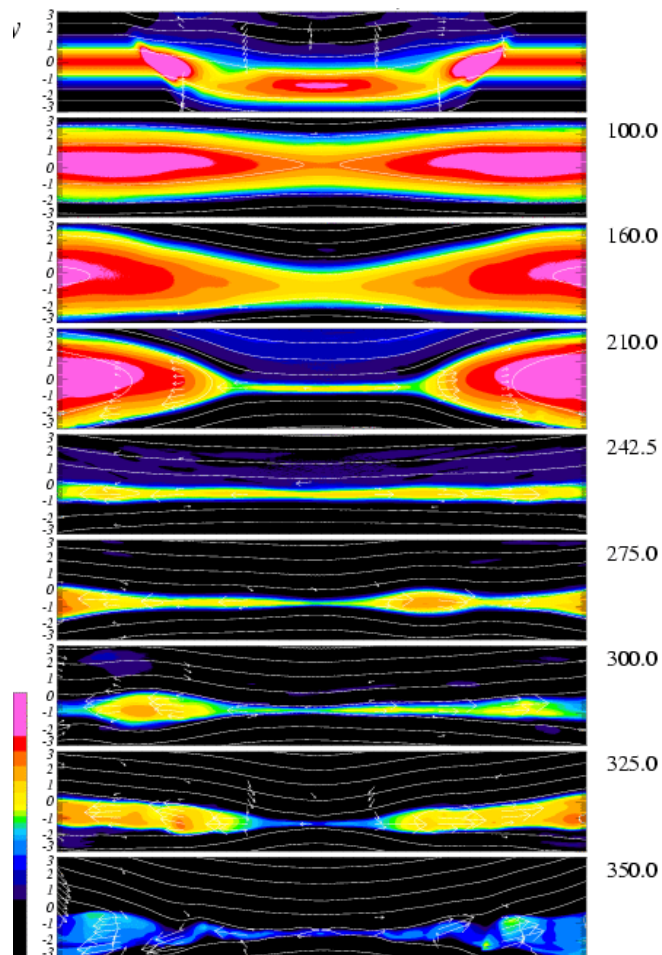
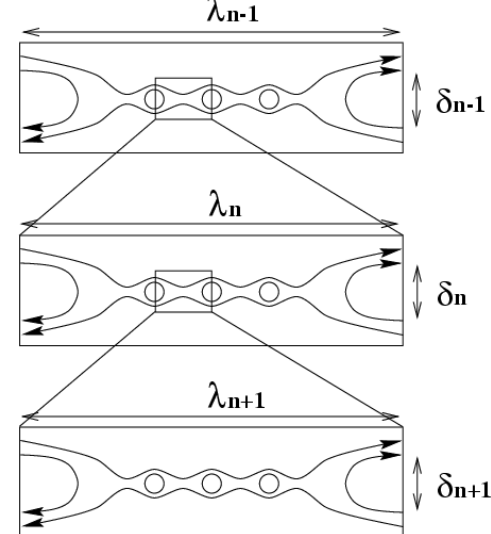
Plasmoid



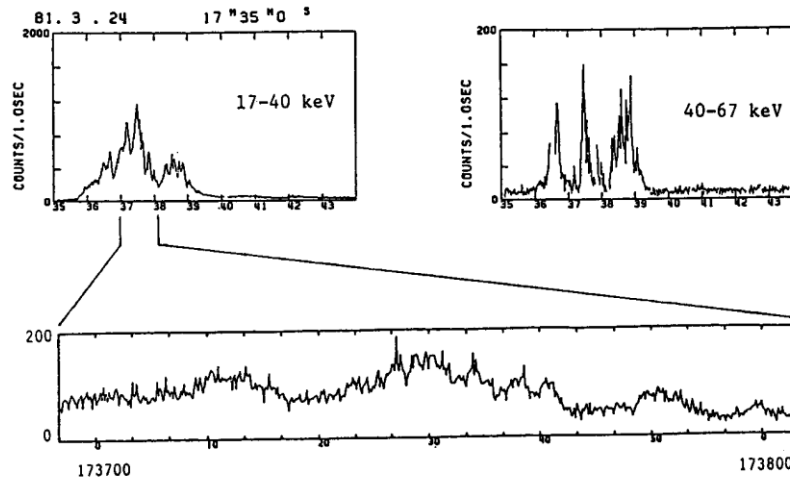
$$\text{Energy release rate} = \frac{dE}{dt} \approx \frac{B^2}{4\pi} V_{in} L^2 \approx 10^{-2} \frac{B^2}{4\pi} V_A L^2$$

# MHD simulations show plasmoid-induced reconnection in a fractal current sheet

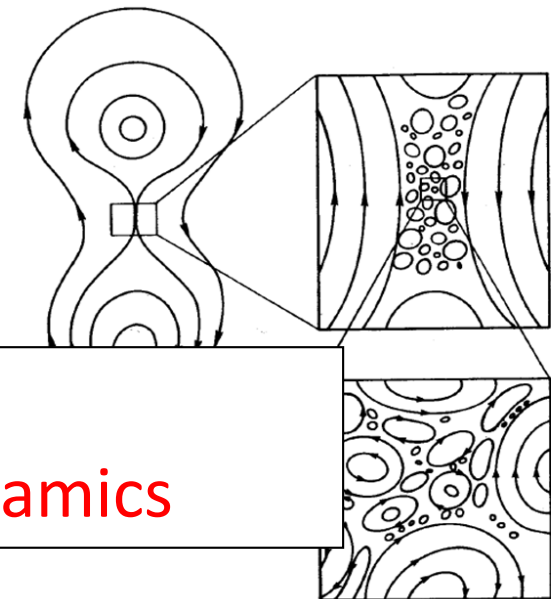
(Tanuma et al. 2001, Shibata and Tanuma 2001)



Observation of hard X-rays and microwave emissions show **fractal-like time variability**, which may be a result of fractal plasmoid ejections



(Tajima-Shibata 1997)



**This fractal structure enable to connect micro and macro scale structures and dynamics**

**Fractal current sheet**

Benz and Aschwanden 1989

Lazarian and Vishniac 1998

Zelenyi 1996, Karlicky 2004, Barta, Buechner et al. 2010

Bhattacharjee et al. 2009, Loureiro et al. 2012

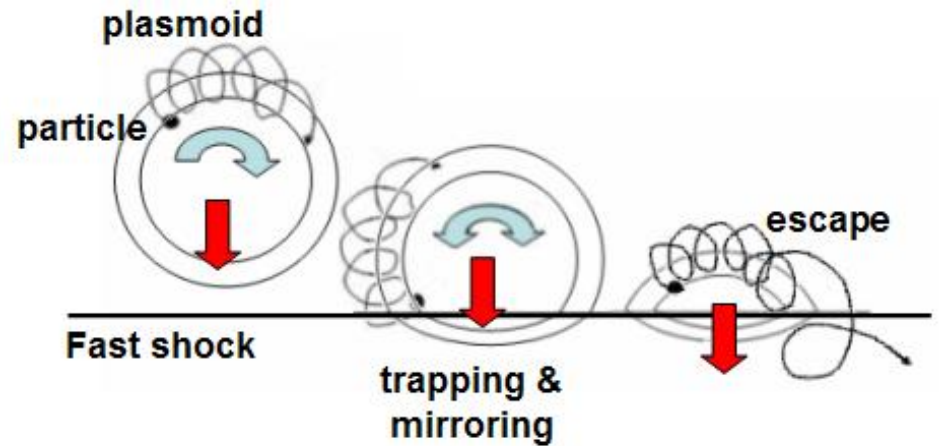
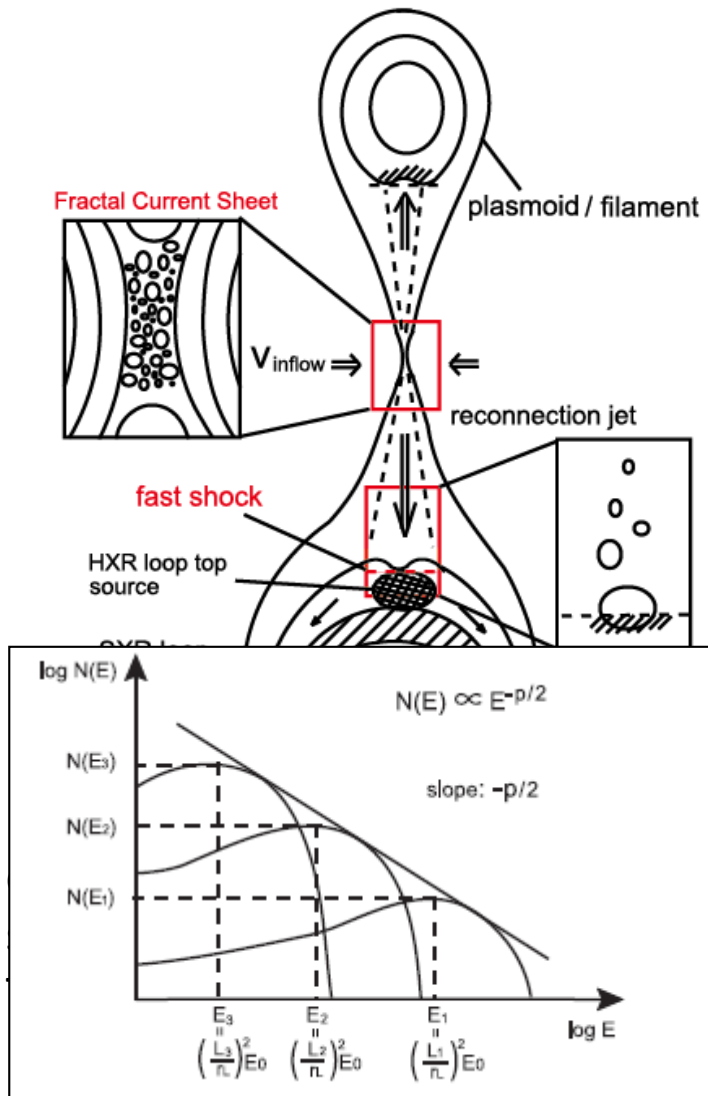


Small-scale electric fields in magnetic X and O points

Aschwanden 2002

# Fractal Reconnection & Particle Acceleration by plasmoids colliding with fast shocks

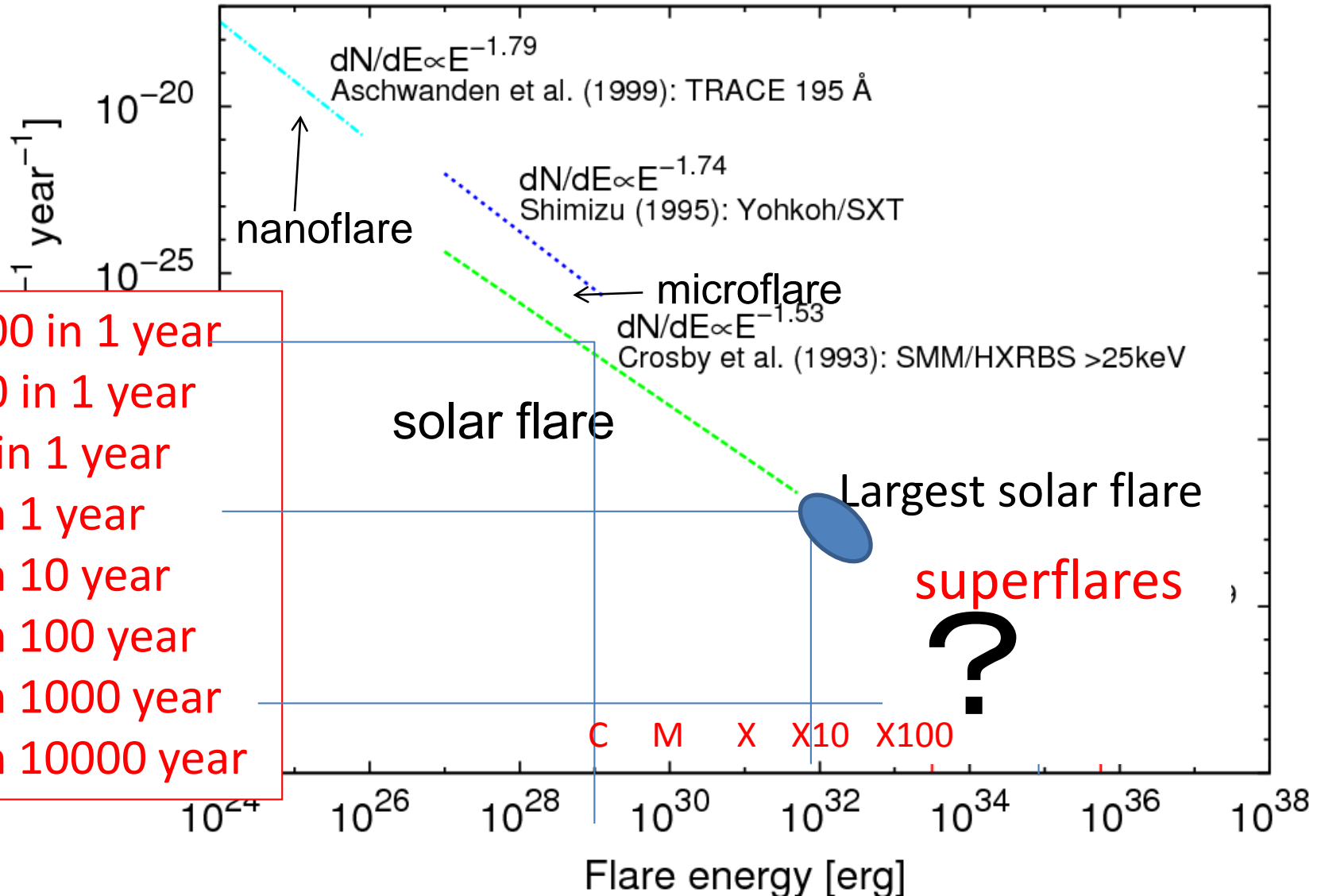
[Nishizuka & Shibata 2013, Phys. Rev. Let. . 110, 051101]



- 1) Particles are **trapped** in a plasmoid.
- 2) Multiple plasmoids collide with fast shock.
- 3) Particles are **reflected** due to **magnetic mirror** effect.
- 4) Reflection length becomes **shorter and shorter**.
- 5) Particles are accelerated by **Fermi process**, until reflection length becomes comparable to ion **Larmor radius**.

# Superflares on Solar Type Stars

# statistics of occurrence frequency of solar flares, microflares, nanoflares



- 1000 in 1 year
- 100 in 1 year
- 10 in 1 year
- 1 in 1 year
- 1 in 10 year
- 1 in 100 year
- 1 in 1000 year
- 1 in 10000 year

# How can we observe superflares on the Sun ?

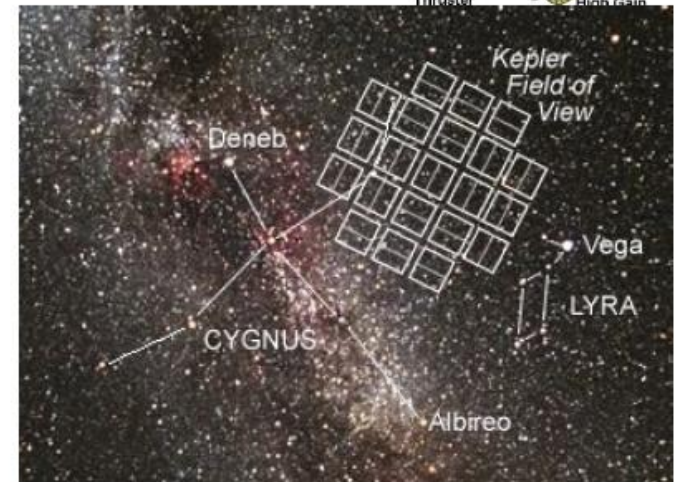
- If empirical statistics rule of solar flares is applied to much larger flares (superflares), then **the frequency of superflares with energy 1000 times larger than the largest solar flares might occur once in 10000 years.**
- However, the period of modern observations of the Sun with telescope is only 400 years.
- How can we observe the Sun for 10000 years ?
- **If we observe 10000 solar type stars (similar to our Sun) for 1 year, we can get the data similar to the data obtained from 10000 years observations of the Sun !**

Prof Sekiguchi kindly told me that the Kepler satellite is taking such data !



# Kepler satellite (NASA)

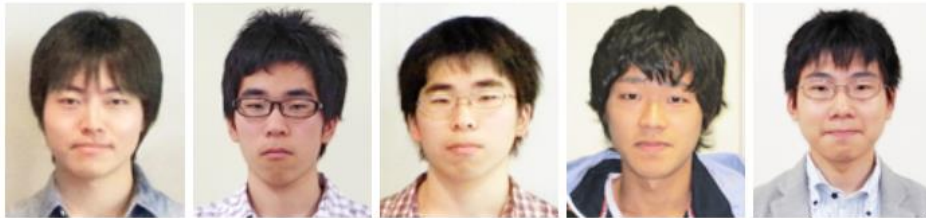
- Space mission to detect exoplanets by observing transit of exoplanets
- 0.95 m telescope
- Observing 160,000 stars continuously (from 2009 to 2013). Among them, 80000 are solar type stars.
- ~30 min time cadence (public data)



# Superflares on Solar Type Stars :

## Our study (Maehara et al. 2012)

- Hence we searched for superflares on solar type stars using Kepler satellite data, which include data of 80000 solar type stars
- Since the data are so large, we asked **1<sup>st</sup> year undergraduate students** to help analyzing these stars,  
because students have a lot of free time (2010 fall)



- Surprisingly, we (they) found **365** superflares on **148** solar type stars (G-type main sequence stars)

# Superflares on solar-type stars

Hiroyuki Maehara<sup>1</sup>, Takuya Shibayama<sup>1</sup>, Shota Notsu<sup>1</sup>, Yuta Notsu<sup>1</sup>, Takashi Nagao<sup>1</sup>, Satoshi Kusaba<sup>1</sup>, Satoshi Honda<sup>1</sup>, Daisaku Nogami<sup>1</sup> & Kazunari Shibata<sup>1</sup>

Undergraduate students

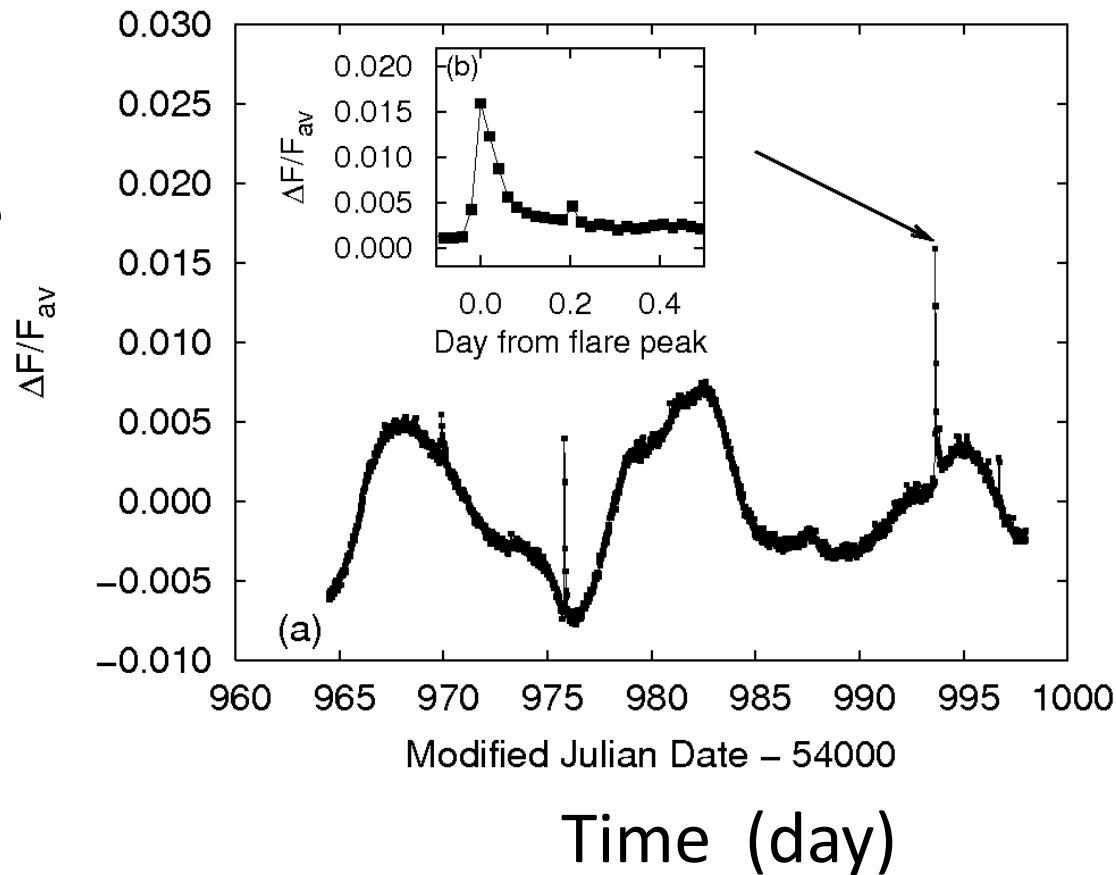
Solar flares are caused by the sudden release of magnetic energy stored near sunspots. They release  $10^{29}$  to  $10^{32}$  ergs of energy on a timescale of hours<sup>1</sup>. Similar flares have been observed on many stars, with larger ‘superflares’ seen on a variety of stars<sup>2,3</sup>, some of which are rapidly rotating<sup>4,5</sup> and some of which are of ordinary solar type<sup>3,6</sup>. The small number of superflares observed on solar-type stars has hitherto precluded a detailed study of them. Here we report observations of 365 superflares, including some from slowly rotating solar-type stars, from about 83,000 stars observed over 120 days. Quasi-periodic brightness modulations observed in the solar-type stars suggest that they have much larger starspots than does the Sun. The maximum energy of the flare is not correlated with the stellar rotation period, but the data suggest that superflares occur more frequently on rapidly rotating stars. It has been proposed that hot Jupiters may be important in the generation of superflares on solar-type stars<sup>7</sup>, but none have been discovered around the stars that we have studied, indicating that hot Jupiters associated with superflares are rare.

We searched for stellar flares on solar-type stars (main-sequence stars) using data collected by NASA’s Kepler<sup>8</sup> during the period from April 2009 to December 2009 (a brief description of the flare search method is described in the legend of Fig. 1 and a detailed description is provided in Supplementary Information). We used the effective temperature ( $T_{\text{eff}}$ ) and the surface gravity ( $\log(g)$ ) available in the Kepler Input Catalog<sup>9</sup> to select solar-type stars. The selection criteria are as follows:  $5,100 \text{ K} \leq T_{\text{eff}} < 6,000 \text{ K}$ ,  $\log(g) \geq 4.0$ . The number of solar-type stars are 9,751 for quarter 0 of the Kepler mission (length of observation period is about 10 d), 75,728 for quarter 1 (90 d), 83,094 for quarter 2 (90 d) and 3,691 for quarter 3 (90 d).

We found 365 superflares (flares with energy  $> 10^{30}$  erg) on 103 solar-type stars (light curves of each flare are shown in Supplementary Fig. 8 and properties of each flare are listed in Supplementary Table 1). The durations of the detected flares are typically a few hours, and their amplitudes are generally 0.1–1% of the stellar luminosity. The bolometric luminosities and bolometric energy of each flare were estimated from the effective temperature in the Kepler Input Catalog.

# typical superflare observed by Kepler

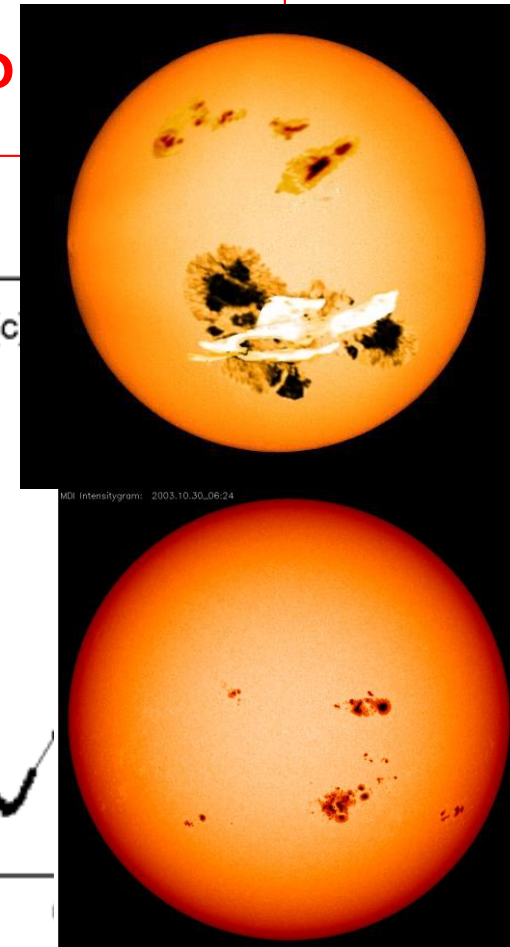
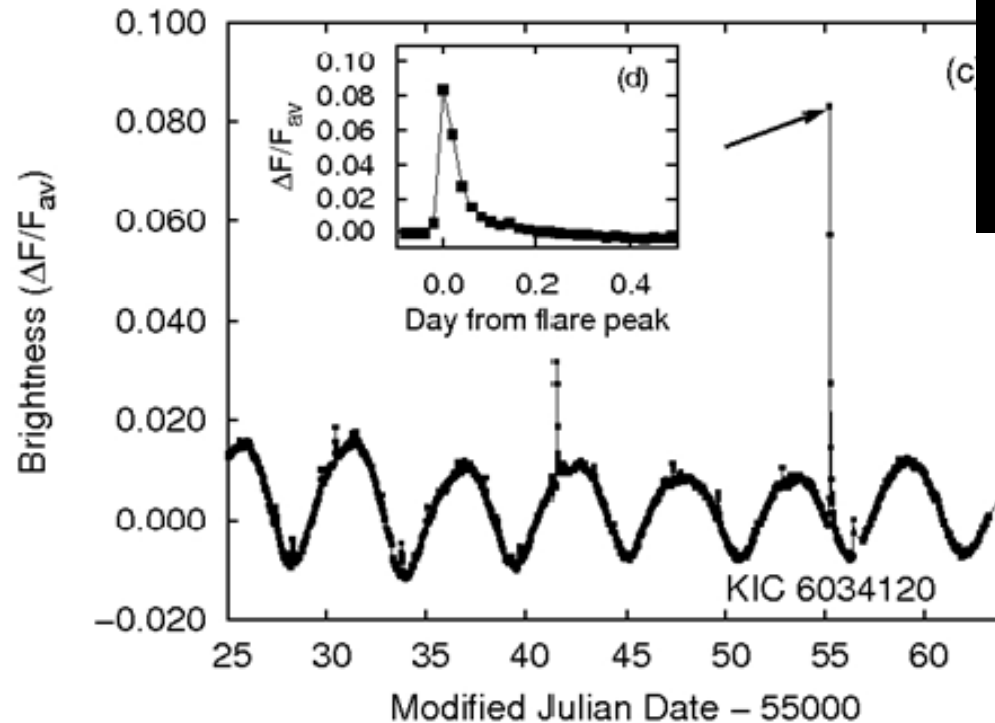
Brightness  
of a star  
and a flare



Total energy  
 $\sim 10^{35}$  erg

# What is the cause of stellar brightness variation ?

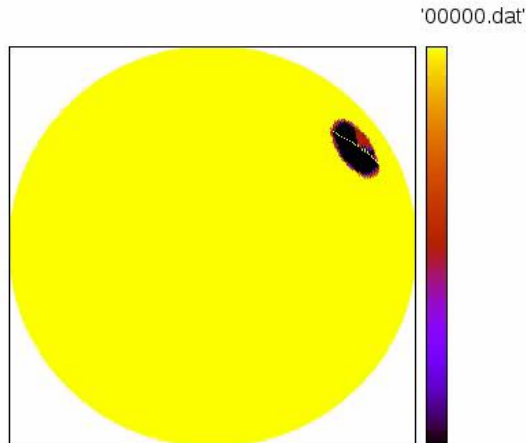
Brightness of a star and a flare



It is likely due to rotation of a star with a big star spot

# Model calculation of stellar brightness variation

**KIC6034120**



**model(green)**

inclination =  $45^\circ$

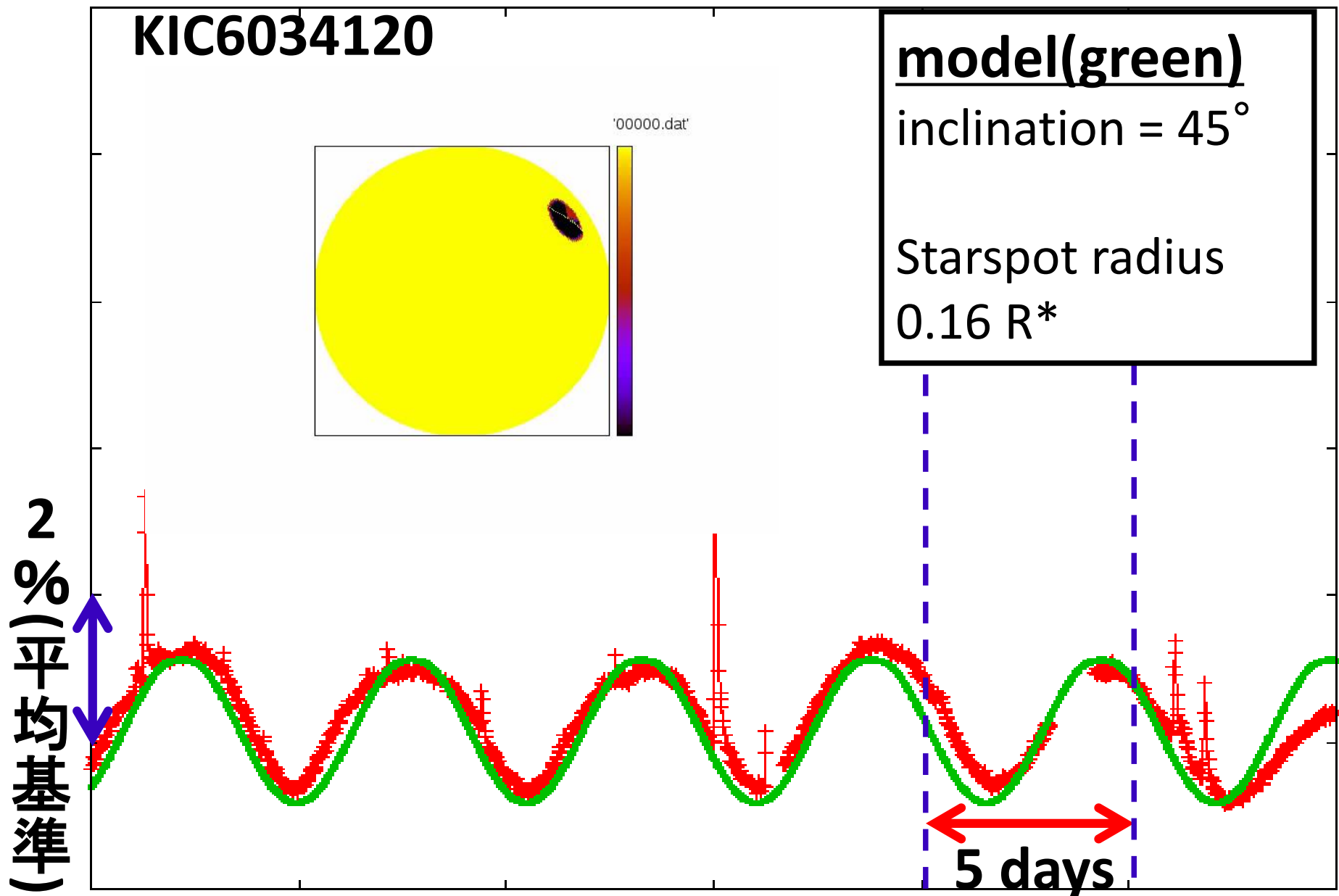
Starspot radius

$0.16 R^*$

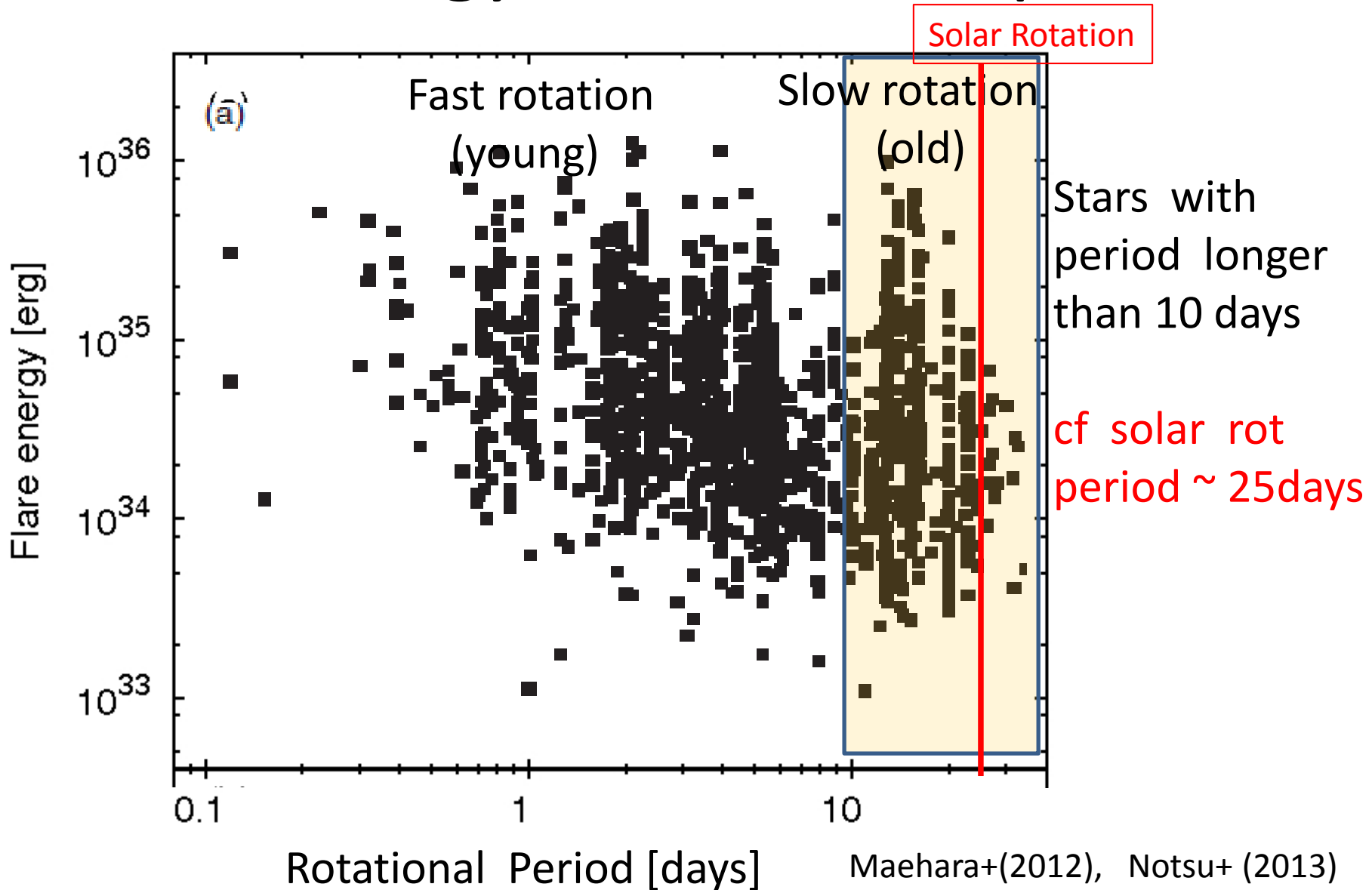
2%  
(平均基準)

5 days

# Model calculation of stellar brightness variation



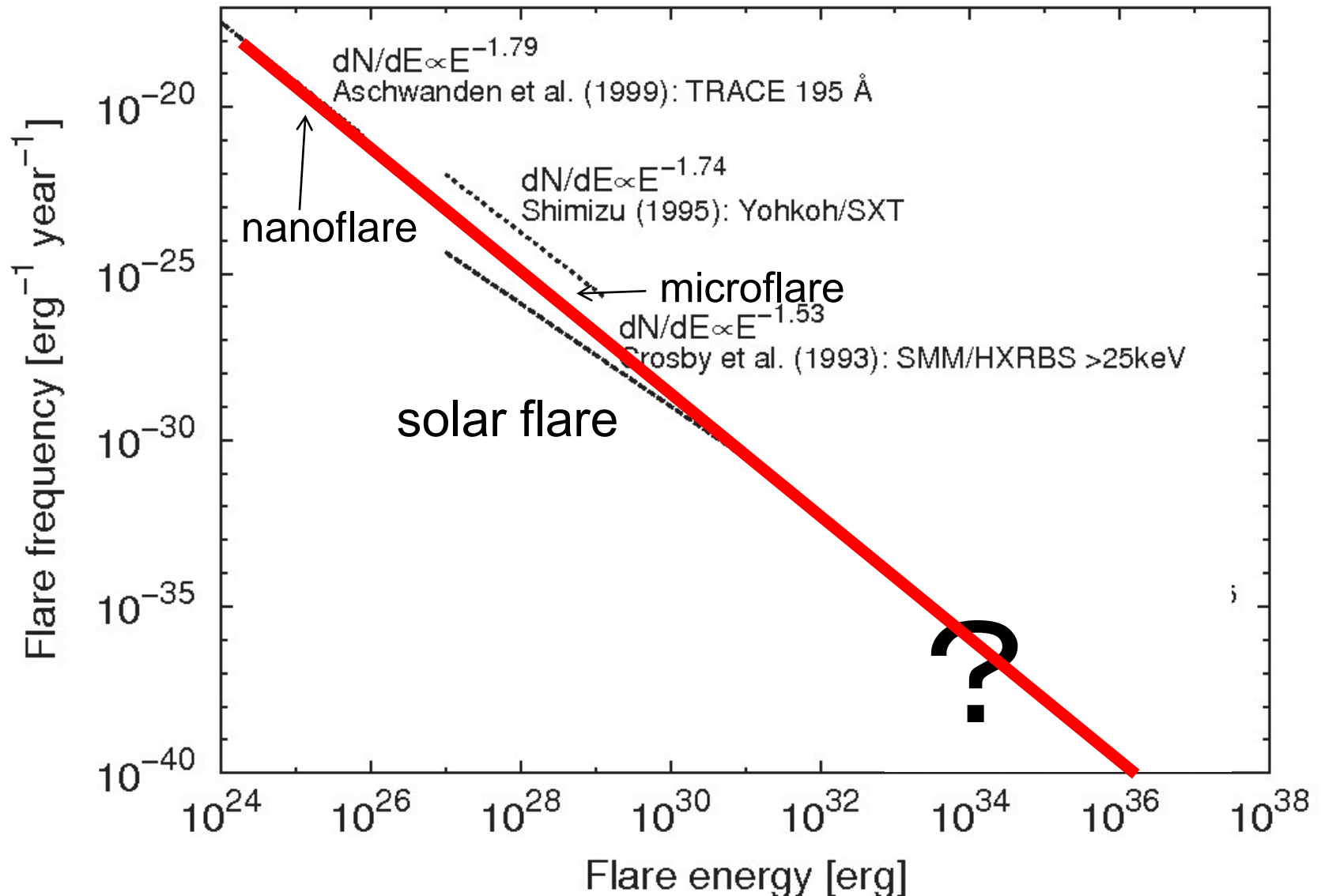
# Flare energy vs rotational period



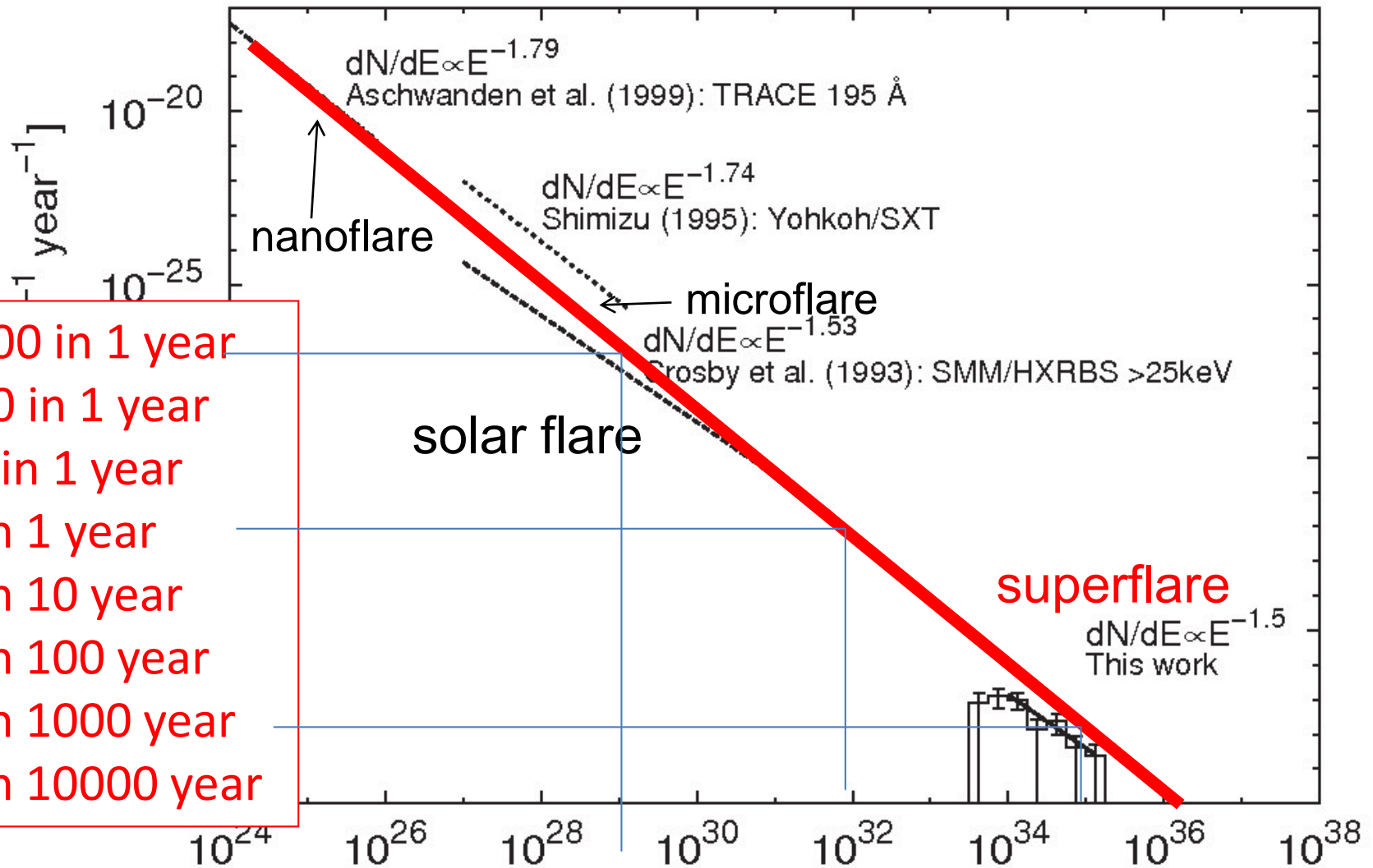


Can Superflares Occur on Our  
Sun ?

# Comparison of statistics between solar flares/microflares and superflares



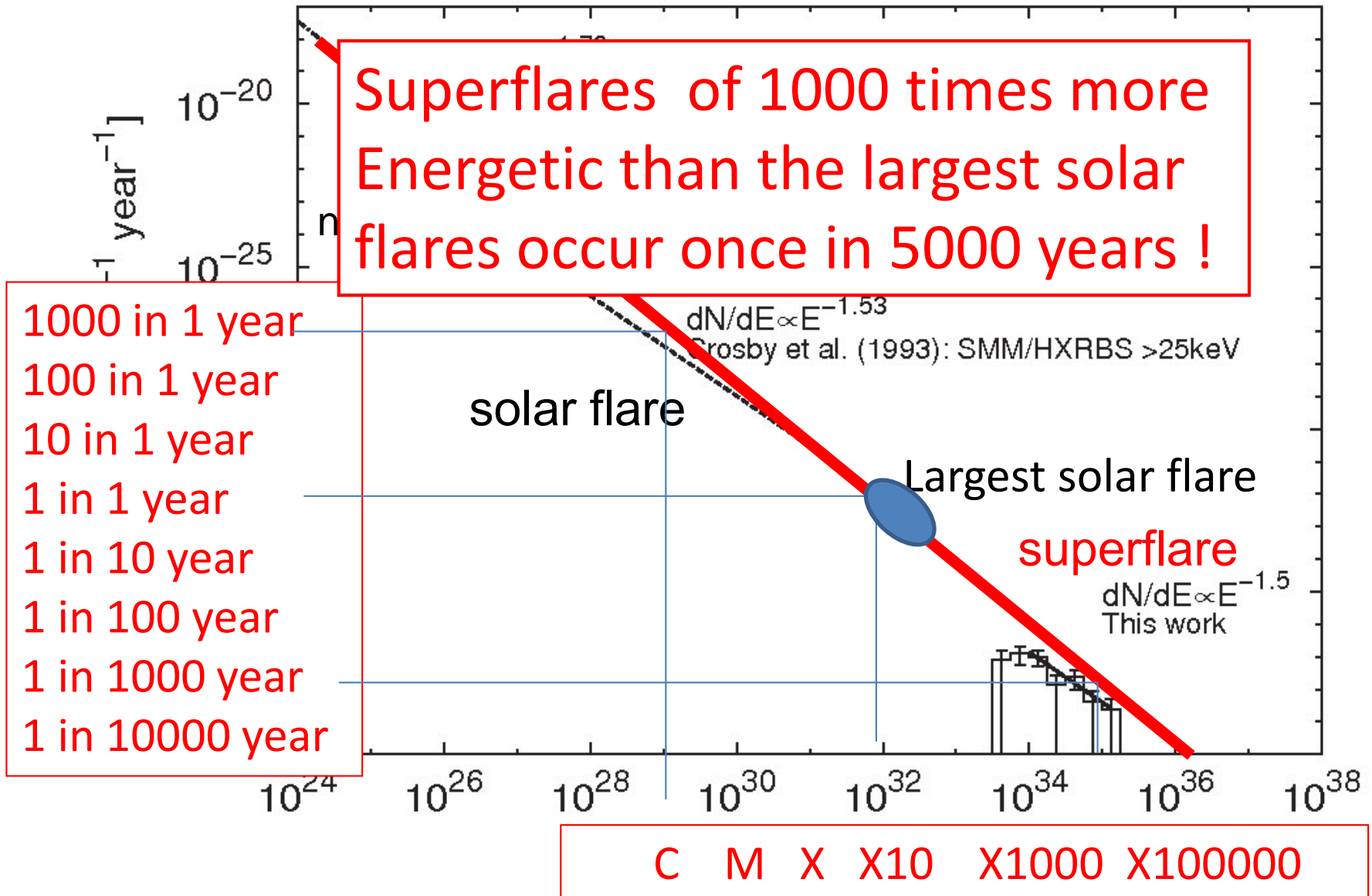
# Comparison of statistics between solar flares/microflares and superflares



1000 in 1 year  
 100 in 1 year  
 10 in 1 year  
 1 in 1 year  
 1 in 10 year  
 1 in 100 year  
 1 in 1000 year  
 1 in 10000 year

C M X X10 X1000 X100000

# Comparison of statistics between solar flares/microflares and superflares



# Evidence of a superflare ?

## LETTER

doi:10.1038/nature11123

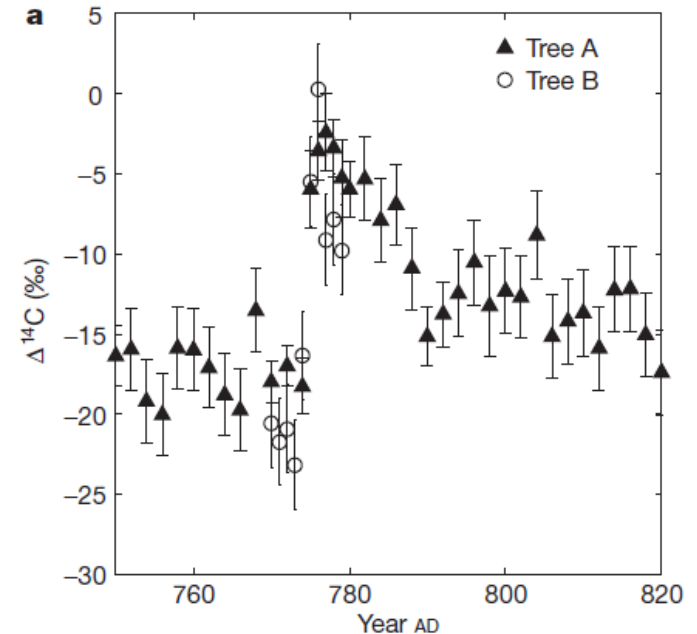
### A signature of cosmic-ray increase in AD 774–775 from tree rings in Japan

Fusa Miyake<sup>1</sup>, Kentaro Nagaya<sup>1</sup>, Kimiaki Masuda<sup>1</sup> & Toshio Naka

Increases in <sup>14</sup>C concentrations in tree rings could be attributed to cosmic-ray events<sup>1–7</sup>, as have increases in <sup>10</sup>Be and nitrate in ice cores<sup>8,9</sup>. The record of the past 3,000 years in the IntCal09 data set<sup>10</sup>, which is a time series at 5-year intervals describing the <sup>14</sup>C content of trees over a period of approximately 10,000 years, shows three periods during which <sup>14</sup>C increased at a rate greater than 3‰ over 10 years. Two of these periods have been measured at high time resolution, but neither showed increases on a timescale of about 1 year (refs 11 and 12). Here we report <sup>14</sup>C measurements in annual rings of Japanese cedar trees from AD 750 to AD 820 (the

Corresponding to 10<sup>34</sup> erg superflare  
If this is due to a solar flare

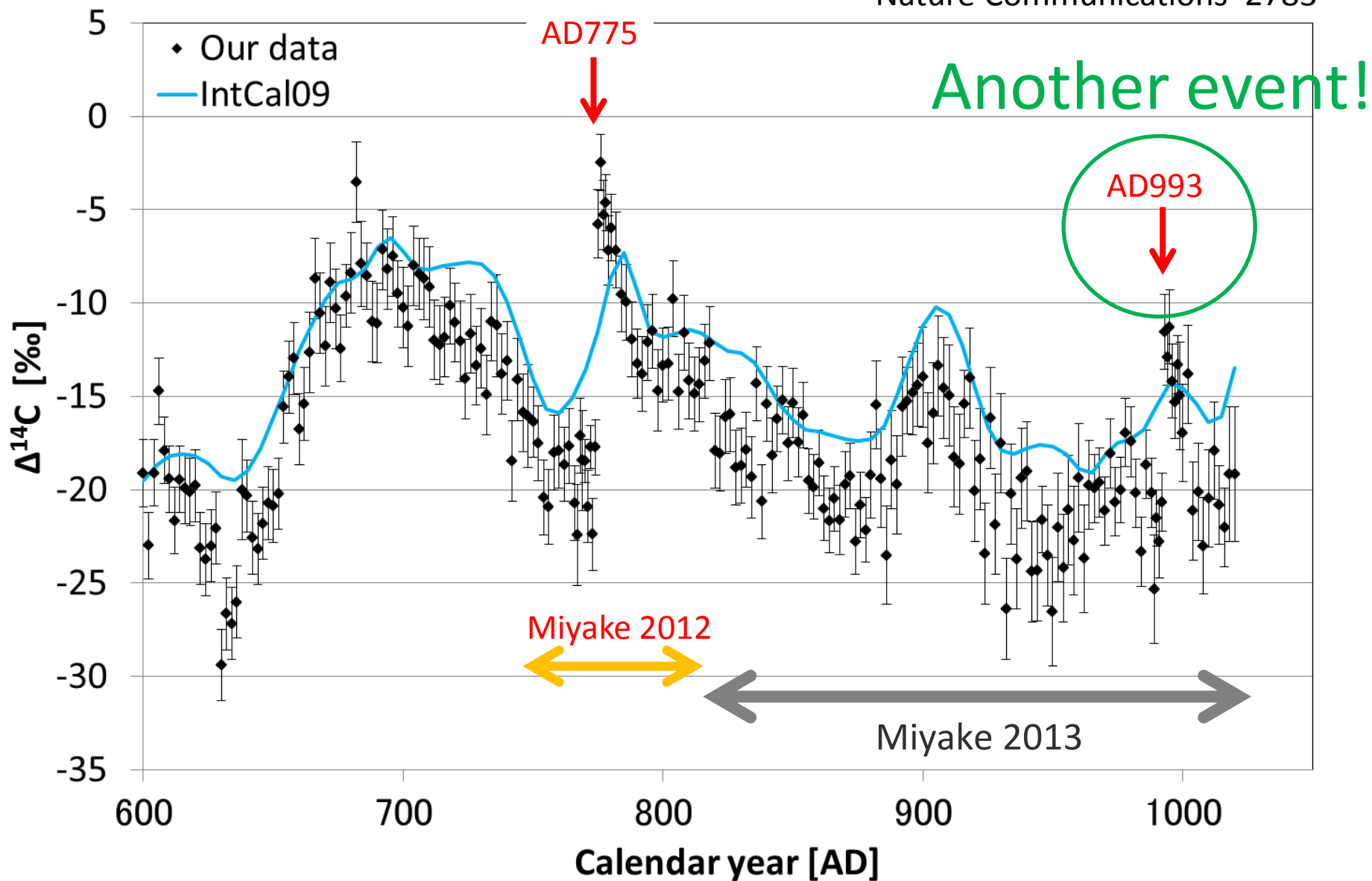
(Miyake et al. Nature ,  
2012, June, 486, 240)



**Figure 1 | Measured radiocarbon content and comparison with IntCal98.** The concentration of <sup>14</sup>C is expressed as  $\Delta^{14}\text{C}$ , which is the deviation (in ‰) of the <sup>14</sup>C/<sup>12</sup>C ratio of a sample with respect to modern carbon (standard sample), after correcting for the age and isotopic fractionation<sup>30</sup>. **a**,  $\Delta^{14}\text{C}$  data for tree A (filled triangles with error bars) and tree B (open circles with error bars) for the period AD 750–820 with 1- or 2-year resolution. The typical precision of a single

# Another evidence ?

From Miyake et al. (2013)  
Nature Communications 2783



If superflares with energy 1000 times larger than the largest solar flares occur on our Sun, what would happen on our Earth and civilization ?

- All artificial satellites would be damaged ?
- All astronauts and some of airline passengers would be exposed to fatal radiation ?
- Ozone layer depletion would occur ?
- Radio communication trouble would occur all over the world ?
- Global blackout would occur on all over the Earth !?
- All nuclear power stations would lose electricity and hence in a state of meltdown ?

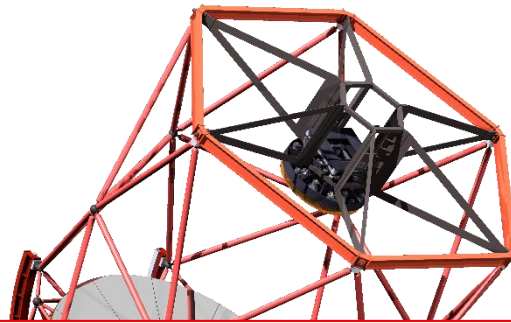
Spectroscopic Observations of Solar type stars causing superflares will be extremely important

## Okayama 3.8m New Technology Telescope of Kyoto Univ (under construction)



### New Technology

1. Making Mirrors with Grinding
2. Segmented mirror
3. Ultra Light mounting



Budget for operation  
Is still lacking.  
Please support us !

High speed photometric and spectroscopic observation of **Transient objects**

Gamma ray bursts  
Exoplanets  
**Stellar flares**  
(superflares)

Will be completed ~ 2017

courtesy of Prof. Nagata (Department of Astronomy , Kyoto University)



# Summary

- Recent observations show **unified view** of solar flares, mass ejections, jets, and nanoflares (Shibata and Magara 2011).
- **Plasmoid-induced-reconnection** and **fractal reconnection** (Shibata and Tanuma 2001) seems to play important role not only **for energy release** but also for **triggering**.
- Kepler data revealed **superflares of  $10^{34}$ - $10^{35}$  erg occur on Sun-like stars** with frequency of **once in 800 - 5000 years** (Maehara et al. 2012). Hence there is a possibility that **superflares of  $10^{34}$  –  $10^{35}$  erg might occur on our present Sun with similar frequency** (Shibata et al. 2013)  
=> dangerous for our civilization !

Thank you for your attention